

Spatial variations in antigorite fabric across a serpentinite subduction channel: Insights from the Ohmachi Seamount, Izu-Bonin frontal arc

Ken-ichi HIRAUCHI¹, Katsuyoshi MICHIBAYASHI², Hayato UEDA³, and Ikuo KATAYAMA¹

¹Department of Earth and Planetary Systems Science, Graduate School of Science, Hiroshima University, 1-3-1 Kagamiyama, Higashi-Hiroshima, Hiroshima 739-8526, Japan

²Institute of Geosciences, Shizuoka University, Shizuoka 422-8529, Japan

³Faculty of Education, Hirosaki University, 1 Bunkyo-cho, Hirosaki, Aomori 036-8560, Japan

Corresponding author: Ken-ichi HIRAUCHI

Phone: +81-82-422-7111

Fax: +81-82-424-0735

E-mail: k-hirauchi@hiroshima-u.ac.jp

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ABSTRACT

We conducted a microstructural study of samples from a natural serpentinite shear zone in the Ohmachi Seamount, Izu-Bonin frontal arc. The serpentinite samples consist mainly of columnar antigorite grains that show marked variations in texture from two approximately orthogonal sets of grains (interpenetrating) to aligned (schistose) forms. Because the two types of grains have similar compositions, these textural differences are interpreted to reflect the existence of a strain gradient toward a plate interface in a subduction zone. The crystal-preferred orientation (CPO) of antigorite with interpenetrating texture is almost randomly oriented, whereas in the case of schistose texture the CPO shows a typical [010](001) pattern. We also found that with increasing intensity of schistosity, the polarization plane of V_{S1} for antigorite grains becomes aligned parallel to the flow plane, consistent with a plane oriented normal to the maximum concentration of slow antigorite c -axes. This configuration results in seismic anisotropy that is approximately five times higher than that for olivine grains. These findings indicate that if a serpentinite layer on the plate interface attains large bulk shear strains ($\gamma > \sim 2$), the resultant alignment of antigorite grains within the layer strongly influences the orientation and magnitude of seismic anisotropy in the mantle wedge, depending on the dip angle of the subducting slab.

1. Introduction

Seismic anisotropy in the Earth's upper mantle is primarily attributed to the deformation-induced preferred orientation of anisotropic minerals such as olivine and orthopyroxene (Christensen, 1984; Mainprice and Silver, 1993; Zhang and Karato, 1995; Zhang et al., 2000). Consequently, the orientation and magnitude of anisotropy is central to understanding the flow geometry and style of deformation in the upper mantle. Measurements of shear-wave splitting, an unambiguous indicator of anisotropy (Long and Silver, 2009), yield the polarization direction of the fast shear wave and the delay time between the fast and slow shear waves (Savage, 1999).

Observations of shear-wave splitting have been carried out for most subduction zones, revealing a wide range of splitting behaviors (Wiens et al., 2008); for example, the direction of fast shear-wave polarization rotates from convergence-normal in areas close to the trench to convergence-parallel farther from the trench (Smith et al., 2001; Nakajima and Hasegawa, 2004; Kneller and van Keken, 2007; Pozgay et al., 2007; Hoernle et al., 2008). Several explanations have been proposed for the development of trench-parallel anisotropy beneath the fore-arc (or arc): deformation of water-rich olivine (Jung and Karato, 2001), aligned melt-pockets (Holtzman et al., 2003), small-scale convection driven by foundering of the arc lower crust (Behn et al., 2007), and trench-parallel flow caused by oblique subduction, slab rollback, or along-strike variations in slab geometry (Smith et al., 2001; Mehl et al., 2003; Kneller and van Keken, 2007). Here, we focus on splitting data collected from local slab earthquakes, as such data constrain the nature of seismic anisotropy in the mantle wedge, whereas the teleseismic ray paths sample the wedge as well as the slab itself and subslab mantle (Long and van der Hilst, 2006).

A large amount of water is liberated from subducting oceanic crust by pressure- and temperature-controlled dehydration reactions, thereby hydrating the overriding mantle wedge and causing serpentinization (Hyndman and Peacock, 2003). Recent seismic observations have documented highly resolved, low-velocity anomalies in the mantle wedge, strongly indicating the presence of serpentine (e.g., Kamiya and Kobayashi, 2000; Bostock et al., 2002; DeShon and Schwartz, 2004; Matsubara et al., 2008). Serpentine minerals (lizardite, chrysotile, and antigorite) are hydrous phyllosilicates (~ 13 wt% H_2O), representing a possibly important water reservoir in subduction zones (Schmidt and Poli, 1998). Lizardite and chrysotile are only stable at temperatures below ~ 300 °C (Evans, 2004), whereas antigorite is stable over a wide range of P - T conditions (up to $T = 600$ °C at $P = 4$ GPa) in subduction

51 zone environments (Ulmer and Trommsdorff, 1995; Bromiley and Pawley, 2003;
52 Komabayashi et al., 2005). Kern et al. (1997) and Watanabe et al. (2007) reported that P-
53 wave and S-wave velocities in foliated antigorite serpentinite propagate slowest in a direction
54 normal to the foliation plane (parallel to the antigorite *c*-axis), and that antigorite serpentinite
55 has a seismic anisotropy approximately five times stronger than that of olivine-dominated
56 rocks. Bezacier et al. (2010) determined the elastic constants of antigorite using Brillouin
57 spectroscopy under ambient conditions, attaining high P-wave and S-wave anisotropies (46%
58 and 66%, respectively) for single-crystal antigorite.

59 Recent deformation experiments under conditions corresponding to the mantle wedge
60 have revealed that the slow *c*-axis of antigorite crystals tends to rotate to an orientation
61 normal to the shear plane during deformation, resulting in seismic anisotropy for an aggregate
62 of antigorite that is an order of magnitude stronger than that for an aggregate of olivine
63 (Katayama et al., 2009). Thus, antigorite in the hydrated forearc mantle may play an
64 important role in controlling the orientation and magnitude of seismic anisotropy (Watanabe
65 et al., 2007; Kneller et al., 2008; Katayama et al., 2009; Mainprice and Ildefonse, 2009;
66 Bezacier et al., 2010; Boudier et al., 2010); however, natural examples of antigorite CPO are
67 largely unknown, with the exception of several local examples (Bezacier et al., 2010; Soda
68 and Takagi, in press), despite the reported occurrence of a coherent body of highly foliated
69 antigorite serpentinites with a mantle wedge origin (Nozaka, 2005; Hirauchi, 2006).

70 In the present paper, we report the results of analyses of peridotites and serpentinites
71 collected from the base of the western wall of the Ohmachi Seamount in the Izu-Bonin frontal
72 arc, where the Pacific plate is subducting beneath the Philippine Sea plate (Figs. 1 and 2). The
73 samples were retrieved using the submersible *Shinkai 6500* of the Japan Agency for Marine-
74 Earth Science and Technology (JAMSTEC). The serpentinites show a wide variety of textures
75 and fabrics associated with abundant antigorite, indicating the occurrence of varied
76 metamorphic and deformation processes in a subduction zone setting. We analyzed antigorite
77 fabrics, mineral composition, and rock texture in order to understand how, after hydration of
78 the original peridotite, the aggregate of antigorite evolved with increasing strain to produce a
79 significant CPO that is consistent with the strong trench-parallel seismic anisotropy in the
80 mantle wedge.

81

2. Tectonic setting

82 The Ohmachi Seamount is located at 29°0′–30′N, 140°35′–55′E, ~20 km east of the
83 volcanic front of the Izu-Bonin–Mariana intra-oceanic arc, the eastern margin of the
84 Philippine Sea plate (Fig. 1a). The Philippine Sea plate consists of active and remnant island
85 arcs and backarc basins (Karig, 1971). The Izu-Bonin–Mariana subduction system, which was
86 initiated in the Early or Middle Eocene as the result of westward subduction of the Pacific
87 plate (Stern and Bloomer, 1992; Bloomer et al., 1995), comprises Eocene to Miocene remnant
88 arcs (Kyushu-Palau Ridge and West Mariana Ridge), post-Miocene backarc basins (Shikoku
89 and Parece Vela basins and the Mariana Trough), and the active Izu-Bonin–Mariana arc.
90 Before post-Miocene backarc opening, the Ohmachi Seamount was located within the Eocene
91 volcanic field, whose remnants are the present-day Izu-Bonin–Mariana forearc and Kyushu-
92 Palau Ridge (Stern et al., 2003).

93 The Ohmachi Seamount consists mainly of Late Eocene to Early Oligocene andesite and
94 Early Miocene turbidites (Nishimura, 1992; Yuasa et al., 1998, 1999), and is terminated to the
95 west by a normal fault of a Quaternary rift system (Fig. 1b). Serpentinite is exposed along the
96 base of the fault scarp, overlain by andesite and turbidites. These stratigraphic relationships
97 suggest that the serpentinite body was exhumed during the early stages of development of the
98 Izu-Bonin–Mariana arc system (Ueda et al., 2004). Niida et al. (2001, 2003) reported that
99 lherzolitic peridotites from the serpentinite body originated from a low degree of melting or
100 melt refertilization of depleted mantle peridotites. Ueda et al. (2004) reported a piece of
101 garnet-zoisite amphibolite float obtained from the northernmost part of the serpentinite body
102 that had been subducted to eclogite-facies conditions ($T = 600\text{--}700\text{ }^{\circ}\text{C}$, $P = \sim 2\text{ GPa}$).

103 In April 2008, the northern and central parts of the N–S trending serpentinite body (Fig.
104 1B) were explored during five dives (6k#1064–1068) made by the submersible *Shinkai 6500*
105 during cruise YK08-05 by the R/V *Yokosuka*, in order to collect rock samples through
106 submersible-based geological observations. For example, as part of dive 6k#1066 (observer:
107 K.H.), we crossed the western flank of the Ohmachi Seamount in water depths of 3500–3200
108 m (Fig. 2), collecting samples at seven stations (S1–S7 in Fig. 2). At 3450 m depth,
109 Quaternary sediments (clays and calcareous silts) are deposited on the sea floor (S1 in Fig. 2).
110 At 3415–3350 m depth, cliff exposures are dominated by schistose antigorite serpentinites
111 that strike N–S and dip steeply to the east (S2–S4, S7 in Fig. 2), and that are associated with
112 minor lizardite/chrysotile serpentinites. Talus deposits, including cobbles of massive
113 antigorite serpentinites, cover the slope at 3310–3275 m depth (S5 in Fig. 2). Above 3240 m
114 depth, volcanic rocks (andesite and dacite) are widely exposed on the western flank of the

115 seamount (S6 in Fig. 2). We also collected a cobble of lizardite/chrysotile serpentinite at 3235
116 m depth.

117 **3. Peridotite and serpentinite from the Ohmachi Seamount**

118 *3.1. Petrography*

119 The ultramafic rocks of the Ohmachi Seamount consist of residual peridotite (lherzolite or
120 harzburgite), dunite, wehrlite, and clinopyroxenite, although they are so strongly altered that it
121 is difficult to estimate their modal (mineral) compositions. In particular, olivine and pyroxene
122 are extensively serpentinized in the residual peridotites collected in the present study;
123 however, Niida et al. (2001) made the preliminary observation that spinel lherzolites from the
124 seamount show a protogranular texture. The lherzolite consists dominantly of coarse olivine
125 grains (up to 1 cm in size) with anhedral orthopyroxene and clinopyroxene (Niida et al., 2001).
126 Most anhedral (vermicular) spinel grains (0.2–2.0 mm) are completely replaced by Cr-
127 magnetite surrounded by a corona of chlorite.

128 The wehrlite and clinopyroxenite, which show a coarse-grained cumulus texture, contain
129 euhedral to subhedral clinopyroxene (diopside) (0.5–3.0 mm in size) with interstitial olivine
130 crystals (up to 4 mm in size) that show a prominent cleavage (Fig. 3A) (= cleavable olivine;
131 Ohara and Ishii, 1998). Altered spinel grains (0.1–1.5 mm in size) are euhedral to subhedral.
132 One dunite sample contains small chromitite pods.

133 In the ultramafic rocks, pargasitic amphibole (edenite), a hydrous phase in the upper
134 mantle (Jenkins, 1983), occurs locally as an accessory phase interstitial to clinopyroxene and
135 olivine. Tremolite occurs as relatively coarse-grained (up to 1 mm in size), acicular to
136 prismatic grains, crosscut by antigorite.

137 Serpentine minerals (lizardite, chrysotile, and antigorite) were identified by laser Raman
138 spectroscopy, based on a comparison with the data reported by Rinaudo et al. (2003) and
139 Auzende et al. (2004). Most olivine and pyroxene grains have been directly replaced by
140 antigorite, a high-temperature serpentine species (Fig. 3A) (Evans, 2004). Antigorite crystals
141 are euhedral to subhedral, and 0.1–1.5 mm in length. Some antigorite-bearing samples show a
142 typical interpenetrating texture (Wicks and Whittaker, 1977) (Fig. 3B), characterized by
143 randomly oriented antigorite blades accompanied by magnetite grains with irregular outlines.
144 This mode of occurrence indicates the irregular replacement of primary minerals under static
145 conditions (Maltman, 1978). Some pyroxenes are completely altered to antigorite, occurring
146 as bastite pseudomorphs (Wicks and Whittaker, 1977).

147 Most of the antigorite serpentinites are strongly foliated, indicative of antigorite schist.
148 The foliation and lineation are defined by aligned antigorite blades, and elongate ribbons and
149 tails of fine-grained aggregates of magnetite that occur around spinel porphyroclasts.
150 Antigorite blades that show interpenetrating texture appear to be crosscut by aligned
151 antigorite blades (herein referred to as schistose antigorite); some of the areas of
152 interpenetrating texture occur as lens-shaped relics aligned parallel to the schistosity (Fig. 3C).
153 Relict olivine grains occur as rigid porphyroclasts surrounded by the antigorite matrix (Fig.
154 3D). Elongate porphyroclasts of bastite are aligned parallel to the schistosity (Fig. 3E), as are
155 needle-shaped grains of secondary clinopyroxene (Fig. 3F). The foliation is overprinted by
156 open microfolds, and a weak crenulation cleavage is developed in the limbs of microfolds
157 (Fig. 3G). The antigorite serpentinites are also characterized by an abundance of chlorite that
158 appears to be overgrown by antigorite (Fig. 3H).

159 Olivine and pyroxene grains that survived the high-temperature, antigorite
160 serpentinization are commonly retrogressively serpentinized to lizardite and chrysotile (Fig.
161 3D, I), a low-temperature serpentine species (Evans, 2004), forming mesh texture (rims and
162 cores) and bastite (Wicks and Whittaker, 1977), respectively. Calcite and chrysotile occur
163 mainly as late-stage fracture fill within the serpentinite samples.

164 3.2. Mineral Chemistry

165 Major-element analyses of minerals were performed using an electron probe
166 microanalyzer (EPMA; JEOL JXA-8200) housed at the Natural Science Center for Basic
167 Research and Development (N-BARD), Hiroshima University, Japan, with settings of 15 kV
168 accelerating voltage, 15 nA beam current, and 2–4 μm beam size. The ZAF method was
169 employed for matrix corrections. Representative analyses are listed in Supplementary Table 1.

170 The forsterite (Fo) and NiO contents of olivine grains vary from 88.2 to 90.7 wt% and
171 from 0.29 to 0.44 wt%, respectively. The Mg# [= $\text{Mg}/(\text{Mg} + \text{Fe}^{2+})$ atomic ratio] and Al_2O_3
172 content of primary clinopyroxene range from 0.913 to 0.927 and from 0.97 to 1.79 wt%,
173 respectively. Secondary clinopyroxene differs from primary clinopyroxene in having higher
174 SiO_2 and CaO contents and lower TiO_2 , Al_2O_3 , Cr_2O_3 , and FeO contents (Supplementary
175 Table 1). The Cr# [= $\text{Cr}/(\text{Cr} + \text{Al})$ atomic ratio] and Mg# of spinels in the residual peridotite
176 range from 0.19 to 0.20 and from 0.66 to 0.69, respectively. The transformation from spinel to
177 Cr-magnetite involved a reduction in Al_2O_3 , Cr_2O_3 , and MgO contents, and an increase in
178 TiO_2 and FeO contents (Supplementary Table 1). All the analyzed amphiboles are calcic

179 amphiboles; some display zoning comprising an edenitic hornblende core and tremolite rim.
180 Tremolite grains differ from the tremolite rims in having higher Al₂O₃, FeO, Na₂O, and K₂O
181 contents, and lower SiO₂ and CaO contents (Supplementary Table 1).

182 The measured serpentine compositions are plotted in Si–Al (Fig. 4a) and Cr–Al (Fig. 4b)
183 binary diagrams, and in a Mg–Al–Si ternary diagram (Fig. 5). Lizardite and chrysotile occur
184 mainly as mesh texture (rims and cores) and bastite. These pseudomorphic serpentines
185 strongly reflect the compositional variations among the parent silicates. Serpentine in both
186 rims and cores is characterized by a high Mg/Si ratio (1.43–1.59) and a low Al₂O₃ content
187 (0.00–0.77 wt%). Serpentines in bastite have a low Mg/Si ratio (1.19–1.37) and a high Al₂O₃
188 content (1.08–2.25 wt%), possibly reflecting the lower Mg and higher Al contents of
189 pyroxene compared with olivine (Supplementary Table 1), and contain minor CaO, Cr₂O₃,
190 and TiO₂.

191 Antigorite has higher SiO₂ contents (39.6–43.0 wt%) and lower H₂O (12.2–14.4 wt%) and
192 MgO (37.1–40.0 wt%) contents than do lizardite and chrysotile, reflecting differences in the
193 structural formulae of these species (Dungan, 1974). The antigorite grains have Mg/Si ratios
194 of 1.35–1.46 and do not inherit the compositions of the parent silicates (Supplementary Table
195 1). The interpenetrating and schistose antigorites have similar compositions and variable
196 Al₂O₃ contents (0.56–4.54 wt%), indicating coupled Al-substitution in tetrahedral and
197 octahedral sites (Fig. 5). With increasing Al₂O₃ content, the composition of antigorite appears
198 to be linearly continuous with that of chlorite in a Mg–Al–Si diagram (Fig. 5). Finally, the
199 Al₂O₃ content of antigorite increases slightly with increasing Cr₂O₃ content (Fig. 4b).

200 3.3. Fabric analysis

201 To examine the development of anisotropy within the antigorite serpentinite, we selected
202 for analysis three representative samples (#1065R006, #1064R012, and #1066R013) with
203 varying degrees of schistosity (massive, transitional, and schistose types, respectively; Fig. 6).
204 We then measured the crystal-preferred orientations (CPOs) of antigorite grains in the three
205 samples from highly polished thin sections using a scanning electron microscope equipped
206 with an electron-backscatter diffraction (SEM-EBSD) system (JEOL JSM6300 with HKL
207 Channel5), housed at the Centre for Instrumental Analysis, Shizuoka University, Japan. Thin
208 sections were made in an arbitrary orientation in the massive type, and cut perpendicular to
209 the foliation and parallel to the lineation (i.e., XZ-sections) in the other two types. We
210 determined the crystal orientations of between 212 and 234 antigorite grains per sample (Fig.

211 7), and visually checked the computerized indexation of the diffraction pattern for each
212 crystal orientation. To quantify the degree of CPO development, we determined the fabric
213 strength and distribution density of the principal crystallographic axes using the J -index and
214 pfJ index (see Ben Ismaïl and Mainprice, 1998; Mainprice et al., 2000; Michibayashi and
215 Mainprice, 2004; Michibayashi et al., 2006, 2009) and the M -index (see Skemer et al., 2005).

216 The massive type contains a random CPO of antigorite grains, with a weak bimodal
217 alignment of [100] and [010] axes and several clusters of [001] axes (Fig. 7A). The two
218 concentrations of [010] axes are separated by 90° , possibly reflecting the orientations of
219 roughly orthogonally oriented antigorite blades (Fig. 6A, B). The fabric intensity (pfj index) is
220 similar among the three crystallographic axes, ranging from 1.29 to 1.39.

221 The schistose type shows a distinct alignment of [001] axes normal to the foliation (Z),
222 [010] axes parallel to the lineation (X), and [100] axes subnormal to the lineation and within
223 the plane of the foliation (Y); i.e., a [010](001) CPO pattern, although both the [100] and
224 [010] axes have a weak girdle distribution within the foliation (XY plane) (Fig. 7C). The
225 fabric intensity (pfj index) of [001] axes ($pfj = 5.59$) in the schistose type is much higher than
226 that of the [100] ($pfj = 2.29$) and [010] ($pfj = 2.99$) axes.

227 The CPO pattern for the transitional type shows scattered [100] axes with a weak
228 alignment parallel to the Y direction, relatively concentrated [010] axes parallel to the
229 lineation (X), and [001] axes with a bimodal alignment, in which one mode is normal to the
230 foliation (Z) (similar to [001] axes in the schistose type) and the other might reflect one of the
231 pre-existing alignments observed in the massive type (Fig. 7B). The fabric intensity (pfj
232 index) of [001] axes ($pfj = 1.40$) is weaker than that of [010] axes ($pfj = 1.63$) but stronger
233 than that of [100] axes ($pfj = 1.27$). The fabric strength (J -index and M -index) of the schistose
234 type ($J = 14.9$; $M = 0.074$) is much higher than that of the massive ($J = 4.38$; $M = 0.012$) and
235 transitional ($J = 4.61$; $M = 0.011$) types, which have similar fabric strengths (Fig. 7).

236 **4. Rock seismic properties**

237 The seismic properties of a rock mass can be computed by averaging the elastic constants
238 of the individual grains in all directions and weighting their contribution based on the modal
239 composition of the aggregate (Mainprice and Humbert, 1994; Mainprice et al., 2000). Given
240 that the three samples selected for analysis (Fig. 6) are composed almost entirely of antigorite,
241 we calculated the seismic properties assuming a composition of 100% antigorite (Fig. 8),
242 using the elastic constants of Bezacier et al. (2010) for antigorite and the Voigt–Reuss–Hill

243 averaging scheme (Mainprice et al., 2000). The P-wave anisotropy (AV_p) was calculated as a
244 percentage using the formula $200(V_{p_{\max}} - V_{p_{\min}})/(V_{p_{\max}} + V_{p_{\min}})$, and the S-wave anisotropy
245 (AV_s) was calculated for a specific propagation direction using the formula $200(V_{s_1} -$
246 $V_{s_2})/(V_{s_1} + V_{s_2})$, where V_{s_1} and V_{s_2} are the fast and slow wave velocities, respectively (e.g.,
247 Pera et al., 2003).

248 The seismic velocities (V_p , V_{s_1} , and V_{s_2}) and seismic anisotropy determined for each rock
249 type are listed in Supplementary Table 2. Values of $V_{p_{\max}}$ range from 7.26 to 8.15 km/s
250 ($V_{p_{\text{mean}}} = 7.56$ km/s), $V_{s_{1\max}}$ from 4.10 to 4.55 km/s ($V_{s_{1\text{mean}}} = 4.25$ km/s), and $V_{s_{2\max}}$ from
251 3.94 to 4.09 km/s ($V_{s_{2\text{mean}}} = 4.01$ km/s). As expected based on the CPO data, the analyzed
252 samples show a wide range of seismic anisotropy: AV_p ranges from 10.4 to 31.3% ($AV_{p_{\text{mean}}} =$
253 17.6%) and AV_s from 8.85 to 35.99% ($AV_{s_{\text{mean}}} = 18.19\%$). Note that AV_p and AV_s calculated
254 from the CPOs of ~200 antigorite grains are similar to those in the bulk rock aggregate
255 (antigorite schist) measured by ultrasonic measurements (Kern et al., 1997; Watanabe et al.,
256 2007).

257 Figure 8 shows stereographic projections of V_p and AV_s data, and the orientation of the
258 polarization plane of the fastest V_s (V_{s_1}). For the massive type, the fastest and slowest V_p
259 directions appear to reflect the weak concentrations of [100] and [001] axes, respectively (Fig.
260 8A). For the transitional type, V_p is fastest parallel to the lineation (X), which coincides with
261 the maximum alignment of [010] axes (Fig. 8B). V_p is slowest subnormal to the foliation (Z);
262 i.e., in a direction between the two concentrations of [001] axes. For the schistose type, V_p is
263 fastest subparallel to the lineation (X) and defines a girdle distribution within the plane of the
264 foliation (XY plane); the slowest direction is normal to the foliation (Z), resulting in an axial
265 symmetry, with the [001] maximum as the symmetry axis (Fig. 8C).

266 The observed variation in polarization anisotropies reflects the measured variations in
267 CPOs and fabric intensities (Fig. 8). Polarization anisotropies for the transitional type show a
268 maximum birefringence (AV_s) for propagation directions at low angles to the plane of the
269 foliation (XY plane), whereas those for the schistose type show a maximum birefringence
270 (AV_s) for propagation directions subparallel to the lineation (X) (Fig. 8B, C). With increasing
271 intensity of schistosity, the orientation of the polarization plane of V_{s_1} becomes parallel to the
272 plane of the foliation (XY plane), coincident with a plane oriented normal to the maximum
273 concentration of [001] axes (Fig. 8).

274

5. Discussion

275 *5.1. Metamorphic and deformation history of serpentinite from the Ohmachi Seamount*

276 The serpentinite samples obtained from the Ohmachi Seamount were possibly derived
277 from the base of the mantle wedge, just above the subducting Pacific plate in an immature
278 intra-oceanic arc system (Ueda et al., 2004). These samples provide evidence for progressive
279 hydration involving a paragenetic sequence from Ca–Al amphibole to tremolite, antigorite,
280 and diopside, which is in contrast to the *P–T* history recorded by amphibole blocks (Ueda et
281 al., 2004, 2009).

282 In the lherzolithic peridotites, pargasitic amphibole is in textural equilibrium with mantle
283 minerals. Based on experimental results (Wallace and Green, 1991; Niida and Green, 1999),
284 the mineral assemblage of the lherzolithic peridotites, when considered in combination with the
285 high Na₂O content (up to 3.5%) of pargasitic amphibole (Supplementary Table 1), indicates
286 formation of the peridotites at pressures of <2 GPa and temperatures of 900–1075 °C,
287 assuming water-saturated conditions. The zoned amphiboles have Na₂O-rich pargasitic cores
288 rimmed by Na₂O-poor tremolite (Supplementary Table 1), suggesting decompressional
289 cooling (Niida and Green, 1999). It is also possible that alteration of clinopyroxene and
290 orthopyroxene in the presence of excess H₂O produced olivine and tremolite at lower
291 temperatures (<800 °C) (Berman, 1986; Bucher and Frey, 2002).

292 Antigorite and diopside may have formed from the hydration of olivine and tremolite in
293 the case that aqueous fluids released by dehydration of the subducting slab were added to the
294 overlying peridotite. Because this hydration event lacks a mineral assemblage of antigorite
295 and brucite, it involved temperatures of 450–575 °C for $P_{\text{total}} = P(\text{H}_2\text{O}) = 1$ GPa (Berman,
296 1986; Bucher and Frey, 2002). Since the Al₂O₃ and Cr₂O₃ contents of antigorite tend to vary
297 sympathetically, the Al (Cr) in antigorite may have been supplied directly from the breakdown
298 of spinel or from chlorite coronas that formed around altered spinel grains.

299 High-temperature serpentinization in the overlying mantle wedge initially occurred under
300 near-static conditions, as indicated by the interpenetrating antigorite blades (Fig. 3B).
301 However, the close compositional relationship between the interpenetrating and schistose
302 antigorite grains indicate that after the formation of antigorite in the mantle wedge, the
303 hydrated region was immediately deformed, given that its viscosity is lower than that of
304 olivine (Hilaireret et al., 2007). Therefore, we suggest that the textural differences among the
305 analyzed antigorite rocks reflect the existence of a strain gradient toward the plate interface.
306 Indeed, the outcrops from which antigorite schist was sampled are located close to

307 amphibolite blocks (Ueda et al., 2004) that represent the remains of ancient subducted oceanic
308 crust.

309 In a subduction zone setting, simple-shear deformation is primarily induced by movement
310 of the subducting slab. We consider that the viscosity contrast between olivine and antigorite
311 (Hilaireret et al., 2007) causes decoupling between the slab and the mantle wedge, thereby
312 impeding asthenospheric flow within the hydrated portion of the mantle wedge (Kneller et al.,
313 2007; Wada et al., 2008; Hilaireret and Reynard, 2009). The resulting return flow within a
314 wedge-shaped, serpentinitized subduction channel, as proposed by Gerya et al. (2002),
315 exhumed the amphibolite blocks within the matrix of the serpentinite, prior to at least the
316 initiation of Early or Middle Eocene arc volcanism (Stern and Bloomer, 1992; Bloomer et al.,
317 1995; Stern et al., 2003). We also note that antigorite schist observed in dive 6k#1066 strikes
318 N–S and dips steeply to the east, subnormal to the surface of the present-day Pacific
319 subducting slab. Since the antigorite schistosity and the slab surface were initially oriented
320 parallel to each other, this finding indicates that the serpentinite layer exhumed in a fossil
321 subduction channel were rotated by later tectonics such as post-Miocene back-arc spreading
322 or arc volcanism, although further analysis is required to determine whether the orientation of
323 the schistosity is consistent throughout the entire serpentinite body.

324 The low-temperature serpentine species appear to have formed from relict olivine and
325 pyroxene grains at temperatures below ~ 300 °C (Evans, 2004), during or after late-stage uplift.
326 Pseudomorphic replacement of primary minerals by lizardite and chrysotile suggests that
327 serpentinitization occurred without concurrent deformation.

328 *5.2. Relationship between texture and seismic properties in the serpentinites*

329 The CPO pattern obtained for schistose antigorite grains is characterized by a strong [010]
330 maximum parallel to the flow direction (lineation) and orientation of the (001) plane parallel
331 to the flow plane (foliation) (Fig. 7C), indicating the dominance of a single slip system:
332 [010](001). In contrast, the CPO pattern for interpenetrating antigorite grains is characterized
333 by a near-random orientation with a weak bimodal alignment of [100] and [010] axes (Fig.
334 7A), reflecting the fact that the antigorite blades are roughly oriented in two contrasting
335 directions (Fig. 6A, B). The CPO pattern obtained for a mixture of interpenetrating and
336 schistose antigorite grains shows that the alignment of [010] axes predated the alignment of
337 the other two axes, thereby representing a relic feature of the preexisting, interpenetrating
338 antigorite grains, particularly the [100] and [001] axes (Fig. 7B). In addition, the fabric

339 strength of transitional type samples is not much stronger than that of the massive type (Fig.
340 7A, B; Fig. 9), meaning that the increased alignment of crystal axes was counterbalanced by
341 the influence of the pre-existing CPO.

342 Both compressional and shear waves show relatively strong anisotropies in the massive
343 type (11.2% and 9.72%, respectively), indicating that even if the hydration of olivine to
344 antigorite occurs under static conditions, a considerable amount of anisotropy develops,
345 although its direction may show no relation to flow geometry. With increasing strain,
346 however, both P-wave and S-wave anisotropies become parallel to the maximum alignment of
347 [010] axes (flow direction), and the polarization plane of V_{S1} becomes aligned parallel to the
348 flow plane (Fig. 8).

349 The maximum P-wave and S-wave anisotropies for the schistose type are 31.3% and
350 35.99%, respectively, approximately three times higher than those for the other two types (Fig.
351 8; Supplementary Table 2), meaning that the orientation and strength of seismic anisotropy in
352 the serpentized layer is largely controlled by the CPO of the schistose type. Although we
353 are unable to directly measure strain from naturally deformed samples, the relationship
354 between shear strain and seismic anisotropy (AV_p and AV_s) in experimentally deformed
355 specimens (Katayama et al., 2009) indicates that large bulk shear strains (at least $\gamma > \sim 2$) are
356 required to produce the significant CPO pattern observed in the schistose type by erasing the
357 earlier CPO pattern observed in the massive type (Fig. 9). Given that a low-viscosity
358 serpentized layer upon a subducting plate interface produces strain localization within the
359 layer and subsequent large bulk shear strains (Hilairret et al., 2007; Wada et al., 2008; Hilairret
360 and Reynard, 2009), the above results indicate that the strongly anisotropic serpentinite layer
361 dominates the base of the hydrated mantle wedge.

362 The CPO patterns obtained for the schistose type in the present study and a natural sample
363 reported by Soda and Takagi (in press) reveal an alignment of a -axes normal to the lineation
364 and within the plane of the foliation (Fig. 7c), whereas a pattern obtained for a natural sample
365 reported by Bezacier et al. (2010) reveals an alignment of a -axes parallel to the flow direction
366 (lineation), indicating that in the case of antigorite, at least two slip systems are activate in
367 nature; i.e., [010](001) and [100](001). We suggest that the above discrepancy in terms of a -
368 axis orientation reflects differences in the physical and chemical environments during
369 deformation. For example, the slip system in olivine is known to change with pressure,
370 temperature, stress, and water content (e.g., Jung and Karato, 2001; Jung et al., 2006, 2009;
371 Katayama and Karato, 2006, 2008). Although the experimentally deformed specimens

372 showed that the antigorite *a*-axis becomes oriented parallel to the flow direction in a
373 laboratory environment at a pressure of 1 GPa and temperature of 300–400°C (Katayama et
374 al., 2009), further deformation experiments performed under different physical and chemical
375 conditions are required to determine the cause of a change in slip system within antigorite.
376 However, because only minor variations in seismic velocity exist between the *a*- and *b*-axes
377 (Mainprice and Ildefonse, 2009; Bezacier et al., 2010), seismic anisotropy produced by the
378 CPO of antigorite could be controlled mainly by the orientation of *c*-axes.

379 Boudier et al. (2010) proposed that antigorite nucleates on particular planar defects in
380 olivine within the host peridotite, resulting in an antigorite schistosity. However, the antigorite
381 CPO in the schistose type shows no geometric relationship with the orientation of planar
382 defects in olivine, because it developed by overprinting the pre-existing antigorite CPO in the
383 massive type.

384 *5.3. Implications for seismic anisotropy in subduction zones*

385 In subduction zone settings, assuming simple corner flow controlled by viscous coupling
386 between the subducting slab and the overriding plate, trench-perpendicular anisotropy beneath
387 the back-arc (or arc) is likely to correspond to A-type (or similar) olivine fabric in the mantle
388 wedge (Hall et al., 2000), in which the fast *a*-axis of olivine crystals is preferentially aligned
389 with the flow direction under water-poor conditions and high temperature/low stress (Zhang
390 and Karato, 1995; Jung et al., 2006). Deformation experiments have shown that under water-
391 rich conditions and low temperature/high stress, olivine deforms by (010)[001] slip that
392 produces a B-type fabric; i.e., the fast olivine *a*-axis becomes oriented perpendicular to the
393 flow direction (Jung and Karato, 2001). Because such conditions (suitable for the
394 development of B-type olivine fabric) may be present in parts of the mantle wedge, transitions
395 in olivine fabric may explain the occurrence of trench-parallel fast directions in the mantle
396 wedge beneath the fore-arc (or arc), as observed at northeastern Japan and Ryukyu, where
397 wedge flow is predicted to be dominated by slab-driven 2D corner flow (Nakajima and
398 Hasegawa, 2004; Mizukami et al., 2004; Kneller et al., 2005; Lassak et al., 2006; Long and
399 van der Hilst, 2006; Tasaka et al., 2008; Katayama, 2009). It has also been proposed that
400 trench-parallel flow, combined with an A-type (or similar) olivine fabric, provides an
401 alternative explanation of trench-parallel anisotropy in the mantle wedge. Long and Silver
402 (2008) proposed that trench-parallel anisotropy in the mantle wedge observed at Tonga is
403 caused by trench-parallel flow, because the high-temperature and low-stress conditions

404 associated with rapid slab rollback (Kincaid and Griffiths, 2003) favor the development of an
405 A-type (or similar) olivine fabric rather than B-type.

406 Despite the fact that the delay time depends on the thickness of the anisotropic layer and
407 its magnitude of anisotropy, it shows marked spatial variations. For example, delay time is
408 $\sim 0.1\text{--}0.2$ s in northeast Japan and $\sim 1\text{--}2$ s at Ryukyu, Izu-Bonin, and Tonga (Smith et al.,
409 2001; Nakajima and Hasegawa, 2004; Anglin and Fouch, 2005; Long and van der Hilst,
410 2006). Katayama (2009) showed that the short delay time recorded in northeast Japan is
411 consistent with the occurrence of a thin anisotropic layer ($\sim 10\text{--}20$ km) caused by the crystal-
412 preferred orientation (CPO) of olivine (B-type fabric); however, the longer delay time
413 observed at Ryukyu, Izu-Bonin, and Tonga cannot be explained in terms of the CPO of
414 olivine, because in such a case the calculated anisotropic layer ($\sim 100\text{--}200$ km) would be
415 thicker than the entire mantle wedge sampled by local S waves (Katayama et al., 2009).

416 Recent deformation experiments revealed that with increasing shear strain, the slow
417 antigorite *c*-axis becomes oriented perpendicular to the flow plane, resulting in seismic
418 anisotropy for antigorite crystals that is an order of magnitude stronger than that for olivine
419 crystals (Katayama et al., 2009). Bezacier et al. (2010) measured the CPO of antigorite
420 crystals in a natural sample from the Escambray massif (central Cuba) using SEM-EBSD,
421 revealing that the seismic anisotropy (36.8% for $AV_{p_{\max}}$ and 50.52% for $AV_{s_{\max}}$) is higher
422 than that in the experimentally deformed samples reported by Katayama et al. (2009) (25.8%
423 for $AV_{p_{\max}}$ and 26.27% for $AV_{s_{\max}}$).

424 For the three samples from the Ohmachi Seamount described in the present study, both P-
425 and S-waves show strong anisotropies (up to 31.3% and 35.99%, respectively) in the
426 schistose type, similar to the values reported for the sample from the Escambray massif (see
427 above); consequently, the large delay times recorded at Ryukyu, Izu-Bonin, and Tonga are
428 consistent with the occurrence of a thin anisotropic layer ($\sim 10\text{--}20$ km) caused by the CPO of
429 antigorite. However, we consider that a thicker anisotropic layer is necessary to achieve large
430 delay times of $\sim 1\text{--}2$ s, because in the serpentinite layer located far from the plate interface,
431 which is expected to record only low bulk strains, the CPO for the massive type might have a
432 strong effect in terms of weakening the significant seismic anisotropy represented by the CPO
433 in the schistose type, which is controlled mainly by flow geometry.

434 As mentioned above, recent seismic observations have revealed that serpentinite is formed
435 through the hydration of peridotite by aqueous fluids expelled upward from the subducting
436 slab (Hyndman and Peacock, 2003). In such regions, the (001) plane of antigorite is

437 preferentially aligned with the plate interface as a result of simple-shear deformation induced
438 by movement of the subducting slab (Katayama et al., 2009). Given that local S-waves
439 propagate in a near-vertical direction in subduction zones with a steeply subducting slab (e.g.,
440 Ryukyu, Izu-Bonin, and Tonga), the significant contribution of the slow antigorite *c*-axis,
441 which is oriented normal to the surface of the slab, results in the polarization direction of the
442 fast shear wave being oriented parallel to the trench axis (Katayama et al., 2009). Furthermore,
443 even if trench-parallel flow is dominant in the mantle wedge, as inferred at Izu-Bonin and
444 Tonga (Anglin and Fouch, 2005; Long and Silver, 2008), the slow antigorite *c*-axis remains
445 oriented normal to the steeply subducting slab, thereby producing a strong trench-parallel
446 anisotropy in the mantle wedge. In contrast, in subduction zones with a shallowly dipping
447 slab (e.g., southwest Japan and Cascadia), a weak trench-parallel anisotropy is observed with
448 delay times of ~0.3–0.5 s (Cassidy and Bostock, 1996; Currie et al., 2001; Salah et al., 2008),
449 despite the fact that a high degree of mantle wedge serpentinization (~70–80%) is suggested
450 by the existence of low-velocity anomalies (Bostock et al., 2002; Ramachandran et al., 2005;
451 Matsubara et al., 2008). This weak anisotropy possibly reflects a small contribution by slow
452 antigorite *c*-axes to near-vertically propagating waves due to the shallow dip of the slab,
453 although the delay time for local S-waves generally depends on ray path length through the
454 mantle wedge, which becomes shorter with decreasing slab dip (Long and Silver, 2008).

455 In summary, the wide range of dip angles of subducting slabs, possibly reflecting the age
456 of the slab, and range in bulk shear strains and thickness of the serpentinite layer may explain
457 the variable strength of trench-parallel seismic anisotropy observed in the mantle wedge.

458 **6. Conclusions**

459 We studied mantle-wedge peridotites and serpentinites collected during five dives
460 (6k#1064–1068) of the submersible *Shinkai 6500* during cruise YK08-05 in 2008 at the base
461 of the western wall of the Ohmachi Seamount in the Izu-Bonin frontal arc. Olivine and
462 pyroxene within the peridotites are directly replaced by antigorite, a high-temperature
463 serpentine species. Some antigorite-rich samples show interpenetrating (massive),
464 orthogonally oriented antigorite grains, although most samples are characterized by schistose,
465 aligned antigorite grains that crosscut the interpenetrating grains. The two sets of grains have
466 similar compositions, indicating that the textural differences reflect the existence of a strain
467 gradient toward the interface of a subducting plate.

468 To understand the development of anisotropy within aggregates of antigorite, we
469 measured the crystal-preferred orientations (CPOs) of antigorite grains in three selected
470 samples that vary in terms of the degree of schistosity (massive, transitional, and schistose
471 types). The CPO for the massive type is almost randomly oriented, whereas that for the
472 schistose type is characterized by a [010] maximum parallel to the flow direction (lineation)
473 and orientation of the (001) plane parallel to the flow plane (foliation). The transitional type
474 has a fabric strength similar to that of the massive type, meaning that the increasing alignment
475 of all crystallographic axes in antigorite, possibly arising from simple shear deformation, was
476 counterbalanced by the effect of the pre-existing CPO.

477 The maximum P-wave and S-wave anisotropies for the schistose type are very high
478 (31.3% and 35.99%, respectively), approximately three times higher than those for the
479 massive type (11.2% and 9.72%, respectively). In addition, the polarization plane of V_{S1} for
480 the schistose type is oriented parallel to the flow plane. These findings indicate that the
481 textural transition of antigorite grains from interpenetrating to schistose forms yields much
482 stronger anisotropy than that of olivine grains, significantly affecting the orientation and
483 magnitude of seismic anisotropy in the mantle wedge, depending on the dip angle of the (001)
484 plane of antigorite. For example, the strong trench-parallel anisotropy recorded at Ryukyu,
485 Izu-Bonin, and Tonga is explained by the occurrence of a thin anisotropic layer caused by the
486 CPO of schistose antigorite grains upon a steeply subducting slab, given the significant
487 contribution to near-horizontally polarized S-waves of the slow antigorite c -axis.

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725 **Captions**

726 Figure 1. (A) Bathymetric map of the Izu-Bonin–Mariana arc–back-arc region. Contour
727 interval is 1000 m. (B) Seafloor topography and inferred geology around the Ohmachi
728 Seamount (1: Quaternary volcanic rocks, 2: basin-fill deposits, 3: Neogene sedimentary rocks,
729 4: Paleogene volcanic rocks, 5: serpentinite, 6: fault scarp). Contour interval is 500 m.
730 Modified from Ueda et al. (2004).

731 Figure 2. (A) Bathymetric map showing the dive track (dark line) of the *Shinkai 6500* (dive
732 6k#1066) and sampling locations (S1–S7). Contour interval is 20 m. (B) Sampling depth
733 versus rock type of recovered samples. Solid and open symbols represent samples collected
734 from outcrop and as float, respectively. atg: antigorite, ctl: chrysotile, liz: lizardite.

735 Figure 3. Optical photomicrographs of serpentinite samples (cross-polarized light). (A)
736 Olivine (ol) crystal with conspicuous cleavage. Note that cleavage planes are infilled with
737 antigorite film. (B) Typical interpenetrating texture, characterized by randomly oriented
738 antigorite (atg) blades. (C) Schistose antigorite (atg) blades with lens-shaped relics of
739 interpenetrating blades. IPT: interpenetrating texture. (D) Relict olivine (ol) grain surrounded
740 by schistose antigorite (atg) blades. Note the occurrence of lizardite (liz) crystals as mesh rim
741 within the olivine grain. (E) Elongate porphyroclast of bastite (BS) aligned parallel to the
742 antigorite (atg) schistosity. (F) Acicular diopside (di) grains aligned parallel to the antigorite
743 (atg) schistosity. (G) Schistose antigorite (atg) blades along with magnetite (mgt) ribbons
744 overprinted by open microfolds, the limbs of which contain a weak crenulation cleavage. sp:
745 spinel. (H) Interpenetrating antigorite (atg) blades that appear to be partially replaced by
746 chlorite (chl). cal: calcite. (I) Altered (serpentinized) olivine grain within schistose antigorite
747 (atg) blades. Note the occurrence of homogeneous mesh texture within the altered grain,
748 comprising polyhedral cores (cr) surrounded by fibrous rims.

749 Figure 4. Compositional trends obtained for serpentine within serpentinite from the Ohmachi
750 Seamount. (A) Si vs. Al. (B) Cr vs. Al. Cations are calculated based on seven oxygens. atg:
751 antigorite, liz: lizardite.

752 Figure 5. Compositional trends obtained for serpentine and chlorite within serpentinite from
753 the Ohmachi Seamount, plotted in a Mg–Al–Si ternary diagram (symbols as in Fig. 4).

754 Figure 6. Optical photomicrographs of three selected samples of antigorite serpentinites from
755 the Ohmachi Seamount (cross-polarized light). Scale bars are all 1 mm. The

756 photomicrographs in (B, D, F) were taken with the gypsum plate inserted. (A and B) Massive
757 type (sample #1065R006). (C and D) Transitional type (sample #1064R012). (E and F)
758 Schistose type (sample #1066R013).

759 Figure 7. Pole figures showing the orientations of the crystallographic axes of antigorite
760 grains. (A) Massive type (sample #1065R006). (B) Transitional type (sample #1064R012).
761 (C) Schistose type (sample #1066R013). Equal-area, lower-hemisphere projections. Contours
762 are in multiples of uniform distribution (m.u.d.). In the case that structural features could be
763 observed in the analyzed samples (i.e., B and C), foliation is oriented vertical in the figure
764 (XY plane; solid line) and lineation (X) is horizontal within the plane of the foliation. J , M ,
765 and pdf are the fabric intensities calculated after Ben Ismail and Mainprice (1998),
766 Michibayashi and Mainprice (2004), and Skemer et al. (2005), respectively.

767 Figure 8. Seismic properties calculated from the CPOs of antigorite within the analyzed
768 samples. (A) Massive type (sample #1065R006). (B) Transitional type (sample #1064R012).
769 (C) Schistose type (sample #1066R013). Equal-area, lower-hemisphere projections. Contours
770 are in multiples of uniform distribution (m.u.d.). Contours for V_p (left-hand column) are in
771 km/s, while those for AV_s (middle column) are in % anisotropy. In the right-hand column, the
772 short lines represent the trace of the V_{s1} polarization plane. In the case that structural features
773 could be observed in the analyzed samples (i.e., B and C), foliation is oriented vertical in the
774 figure (XY plane; solid line) and lineation (X) is horizontal within the plane of the foliation.

775 Figure 9. Relationship between seismic anisotropy (AV_p and AV_s) and fabric strength (M -
776 index) for aggregates of antigorite in natural samples (present study; black symbols) and in
777 experimentally deformed samples (Katayama et al., 2009; gray symbols). SM: starting
778 material, M: massive type, T: transitional type, S: schistose type. Numbers next to data points
779 indicate the magnitude of shear strain (γ) during deformation.

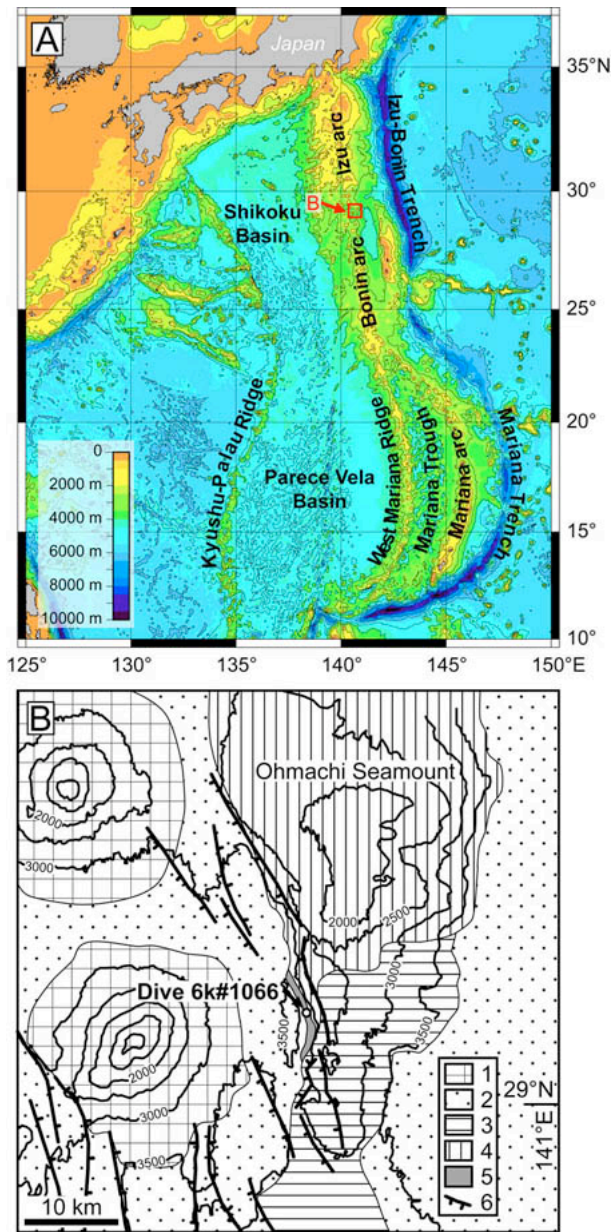


Figure 1

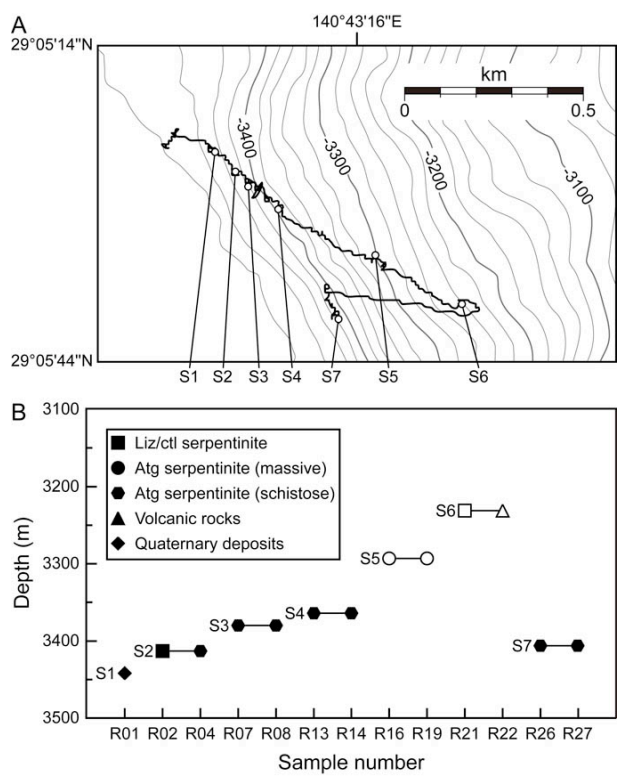


Figure 2

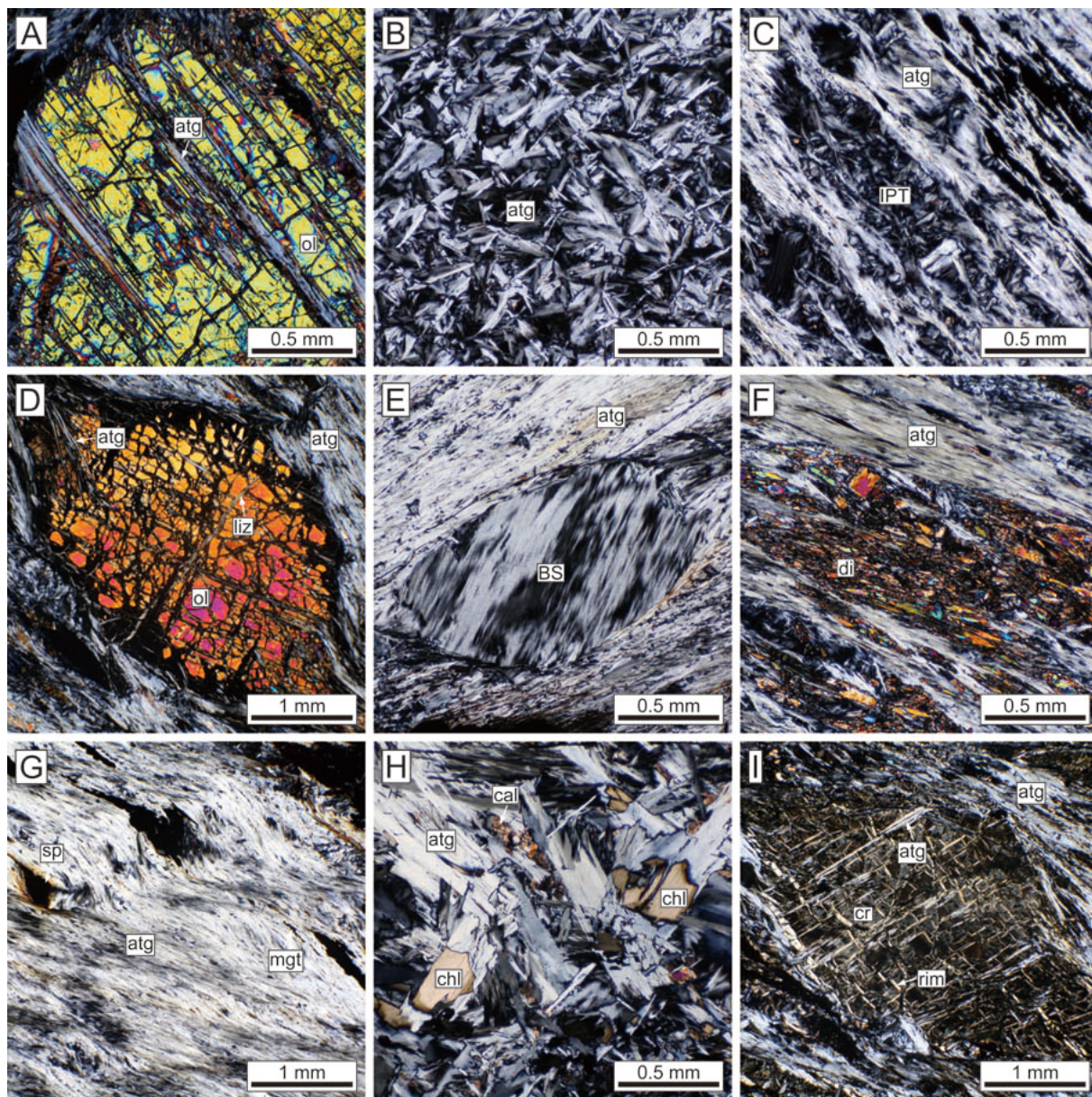


Figure 3

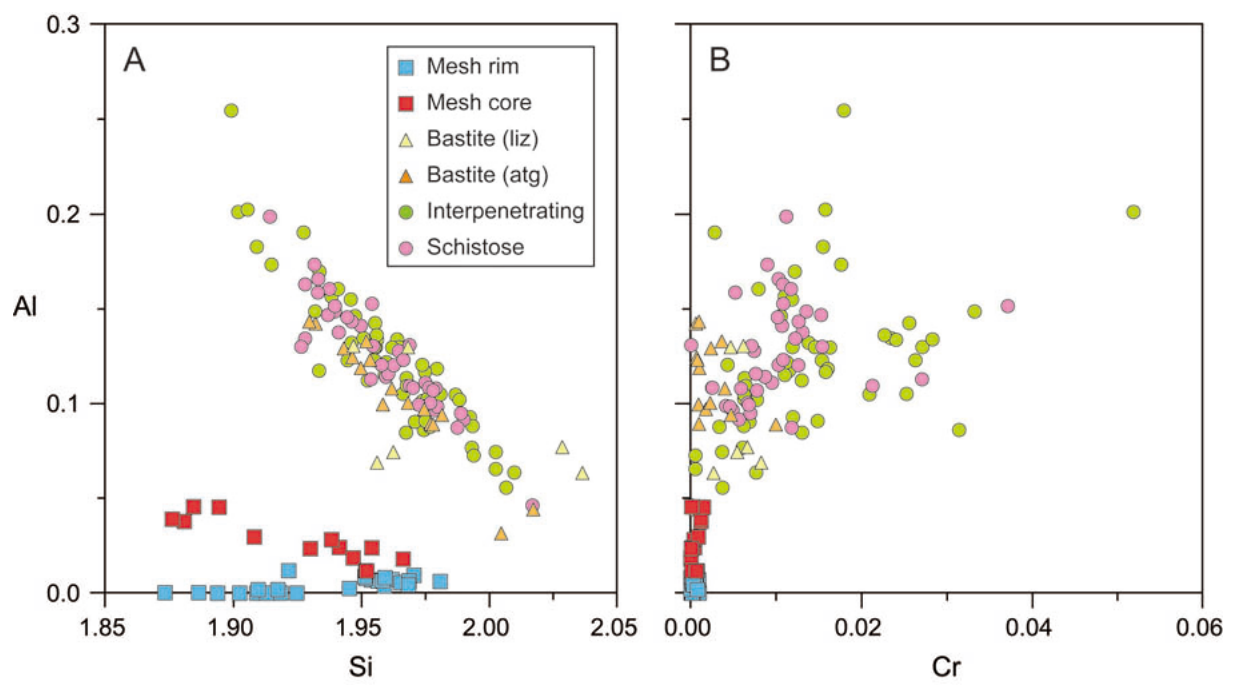


Figure 4

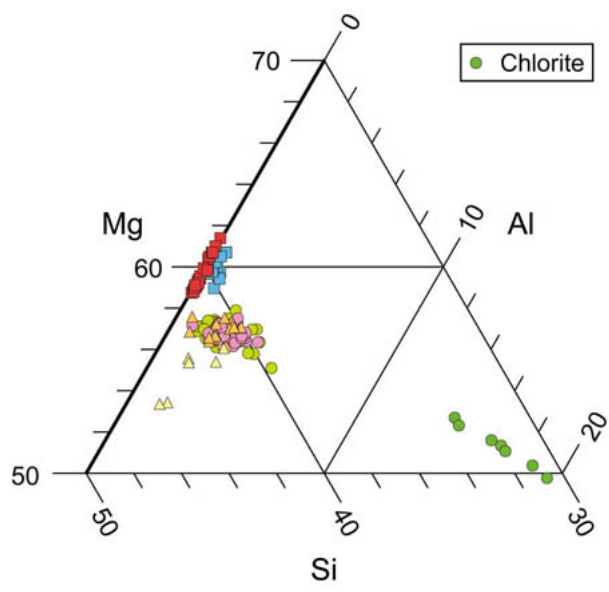


Figure 5

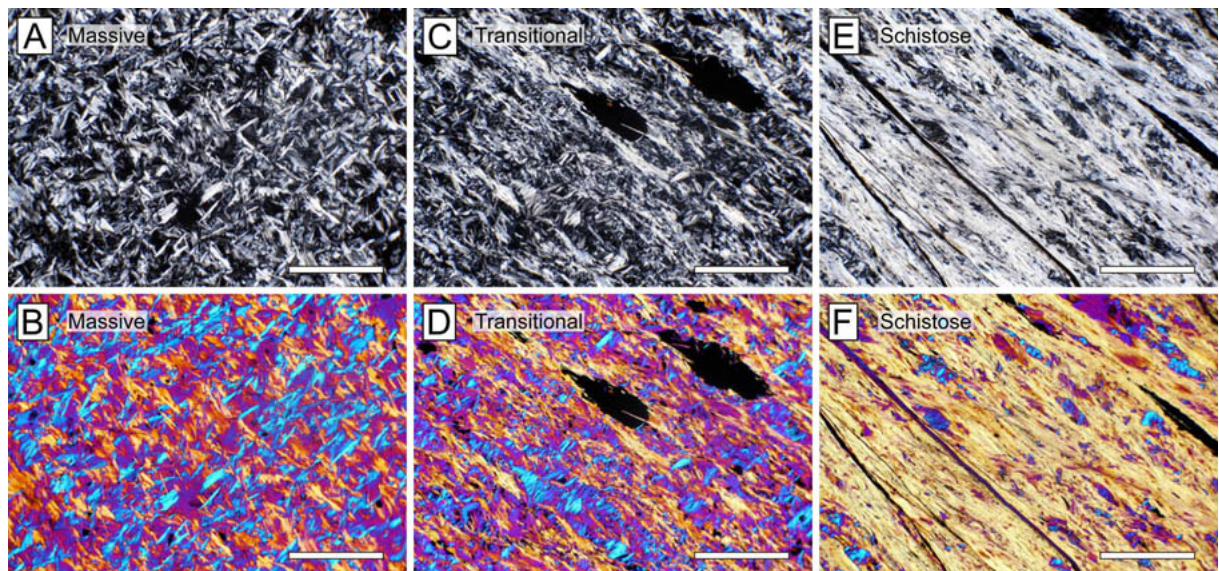


Figure 6

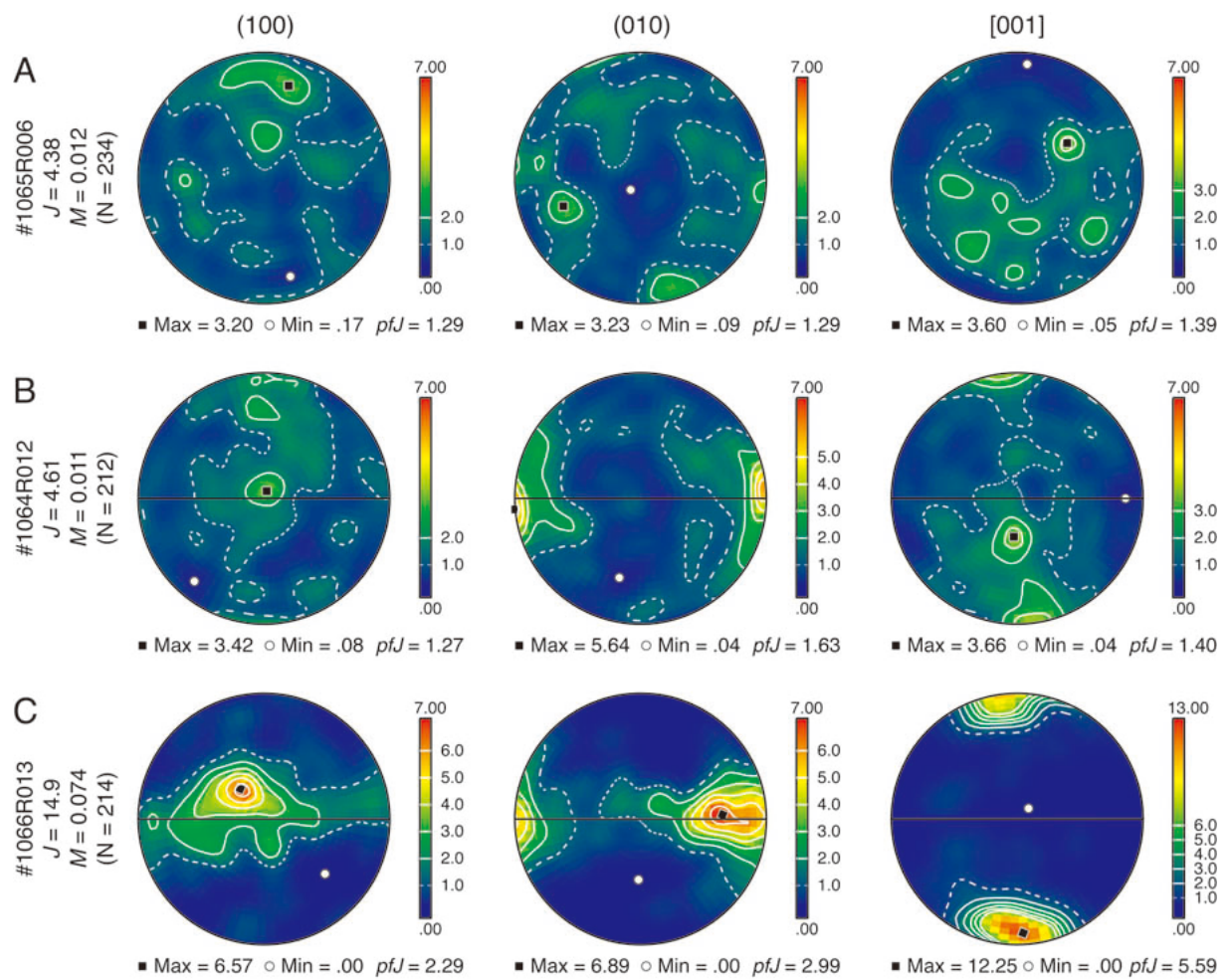


Figure 7

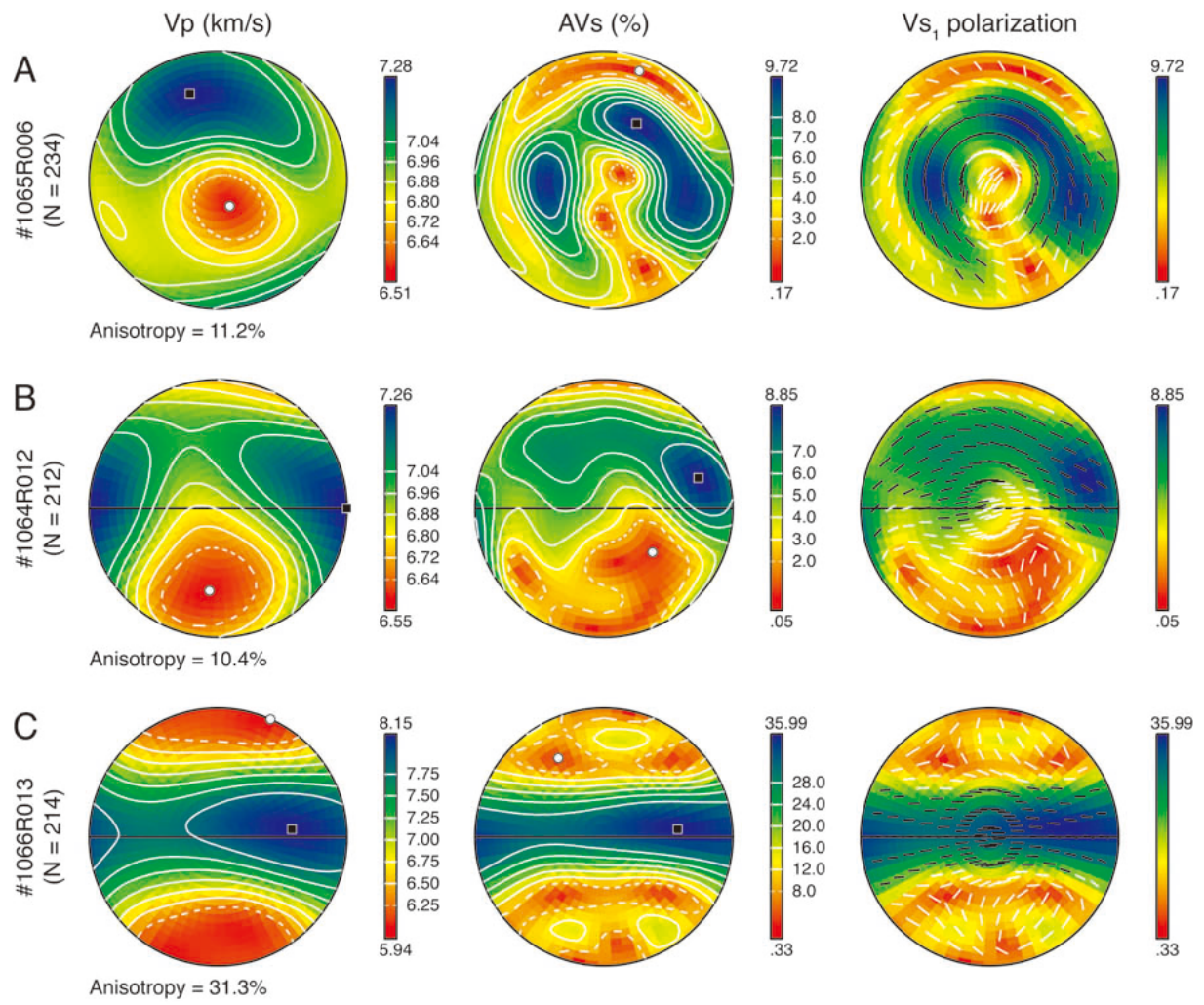


Figure 8

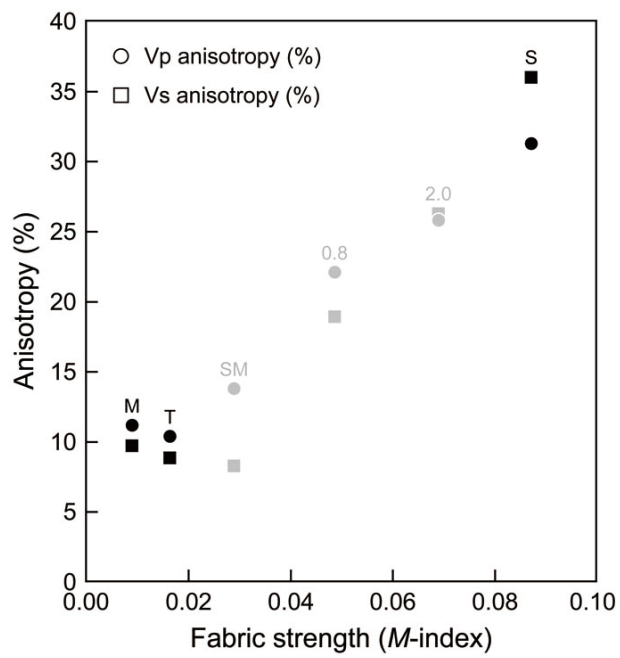


Figure 9

Supplementary Table 1. Representative microprobe analyses of minerals in peridotites and serpentinites from the Ohmachi Seamount

Sample no:	#1064R002		#1065R006		#1066R008		#1066R013			#1066R021			#1066R025			#1068R001			#1068R004		#1068R007		#1068R015			
Rock type	clinopyroxenite		antigorite serpentinite		antigorite schist		antigorite schist			lherzolite			antigorite schist			antigorite schist			antigorite schist		antigorite schist		wehrlite			
Mineral	P-cpx	atg_B	atg_I	atg_I	atg_S	Cr-mgt	sp	liz_MR	liz_B	ol	M-cpx	atg_I	atg_S	chl_I	ol	M-cpx	liz_MR	ctl_MC	M-cpx	ed_core	tr_rim	tr				
SiO ₂	53.3	41.1	42.0	42.0	42.1	0.0	0.0	39.7	39.1	39.8	55.0	41.8	41.6	33.6	41.0	54.9	39.1	39.3	54.6	46.2	57.9	56.0				
TiO ₂	0.09	0.00	0.00	0.01	0.02	0.69	0.10	0.02	0.05	0.00	0.04	0.02	0.01	0.01	0.01	0.02	0.02	0.01	0.00	0.81	0.01	0.06				
Al ₂ O ₃	1.55	2.24	1.58	2.22	1.75	0.02	46.8	0.12	1.26	0.00	0.05	2.83	2.76	12.5	0.02	0.02	0.03	0.41	0.07	10.65	0.24	1.19				
FeO*	2.78	3.23	3.43	3.04	2.99	84.3	18.7	4.56	9.17	11.4	1.61	3.13	3.17	3.52	9.74	1.43	6.80	2.46	1.34	3.91	1.45	1.85				
MnO	0.13	0.07	0.09	0.09	0.05	0.40	0.19	0.04	0.07	0.32	0.01	0.05	0.05	0.03	0.18	0.04	0.14	0.03	0.01	0.08	0.06	0.04				
Cr ₂ O ₃	0.52	0.02	0.09	0.30	0.13	6.76	17.5	0.02	0.14	0.03	0.17	0.32	0.29	0.79	0.00	0.20	0.00	0.00	0.05	0.73	0.00	0.04				
MgO	17.2	39.5	39.4	39.6	39.5	0.9	16.7	39.2	34.0	48.3	17.5	39.3	38.8	35.6	49.9	17.8	39.3	40.9	17.9	18.6	23.6	23.2				
CaO	24.5	0.01	0.00	0.01	0.01	0.01	0.00	0.05	0.13	0.00	25.3	0.01	0.00	0.02	0.01	25.4	0.02	0.02	25.5	12.6	13.0	12.5				
Na ₂ O	0.11	0.01	0.02	0.00	0.00	0.03	0.05	0.03	0.05	0.01	0.41	0.03	0.03	0.07	0.01	0.19	0.00	0.02	0.11	3.23	0.68	1.77				
K ₂ O	0.00	0.00	0.01	0.00	0.01	0.00	0.01	0.00	0.02	0.00	0.00	0.01	0.00	0.03	0.00	0.00	0.01	0.02	0.00	0.38	0.03	0.11				
NiO	0.02	0.07	0.08	0.05	0.08	0.27	0.32	0.29	0.29	0.31	0.00	0.04	0.04	0.02	0.41	0.03	0.16	0.14	0.03	0.07	0.06	0.07				
Total	100.3	86.2	86.7	87.3	86.7	93.4	100.3	84.0	84.2	100.2	100.1	87.6	86.7	86.3	101.3	100.1	85.6	83.3	99.6	97.2	97.0	96.9				
<i>Cations</i>	<i>O = 6</i>	<i>O = 7</i>	<i>O = 7</i>	<i>O = 7</i>	<i>O = 7</i>	<i>O = 4</i>	<i>O = 4</i>	<i>O = 7</i>	<i>O = 7</i>	<i>O = 4</i>	<i>O = 6</i>	<i>O = 7</i>	<i>O = 7</i>	<i>O = 7</i>	<i>O = 4</i>	<i>O = 6</i>	<i>O = 7</i>	<i>O = 7</i>	<i>O = 6</i>	<i>O = 23</i>	<i>O = 23</i>	<i>O = 23</i>				
Si	1.94	1.95	1.98	1.96	1.98	0.00	0.00	1.95	1.96	0.98	2.00	1.95	1.95	1.61	0.99	1.99	1.92	1.93	1.99	6.58	7.96	7.77				
Ti	0.00	0.00	0.00	0.00	0.00	0.03	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.09	0.00	0.01				
Al	0.07	0.12	0.09	0.12	0.10	0.00	1.52	0.01	0.07	0.00	0.00	0.16	0.15	0.71	0.00	0.00	0.00	0.02	0.00	1.79	0.04	0.20				
Fe ²⁺	0.08	0.13	0.14	0.12	0.12	1.19	0.32	0.19	0.38	0.24	0.05	0.12	0.12	0.14	0.20	0.04	0.28	0.10	0.04	0.47	0.17	0.21				
Fe ³⁺						2.24	0.11																			
Mn	0.00	0.00	0.00	0.00	0.00	0.02	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.01	0.01	0.00				
Cr	0.01	0.00	0.00	0.01	0.00	0.26	0.38	0.00	0.01	0.00	0.00	0.01	0.01	0.03	0.00	0.01	0.00	0.00	0.00	0.08	0.00	0.00				
Mg	0.93	2.78	2.77	2.76	2.77	0.07	0.69	2.87	2.54	1.78	0.94	2.73	2.72	2.54	1.80	0.96	2.87	2.99	0.97	3.96	4.83	4.81				
Ca	0.96	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.98	0.00	0.00	0.00	0.00	0.99	0.00	0.00	0.99	1.92	1.92	1.85				
Na	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.03	0.00	0.00	0.01	0.00	0.01	0.00	0.00	0.01	0.89	0.18	0.48				
K	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.07	0.01	0.02				
Ni	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.01	0.00	0.01	0.01	0.00	0.01	0.01	0.01				
Total	4.02	4.99	4.98	4.97	4.97	3.84	3.05	5.04	5.00	3.02	4.01	4.97	4.97	5.03	3.01	4.01	5.08	5.06	4.01	15.87	15.11	15.37				

*Total iron given as FeO; Fe²⁺ and Fe³⁺ of spinel and Cr-magnetite were calculated assuming spinel stoichiometry.

B–bastite, I–interpenetrating, M–metamorphic, MC–mesh core, MR–mesh rim, P–primary, S–schistose.

Supplementary Table 2. Seismic properties (Vp, Vs) of antigorite serpentinites from the Ohmachi Seamount

		Sample no:	Vp			AVs	Vs ₁			Vs ₂			dVs
			Max	Min	Anisotropy	Max	Max	Min	Anisotropy	Max	Min	Anisotropy	Max
			(km/s)	(km/s)	(%)	(%)	(km/s)	(km/s)	(%)	(km/s)	(km/s)	(%)	(%)
Hill average	A	#1065R006	7.28	6.51	11.2	9.72	4.10	3.88	5.3	4.01	3.63	10.0	0.38
	B	#1064R012	7.26	6.55	10.4	8.85	4.11	3.83	7.0	3.94	3.76	4.6	0.35
	C	#1066R013	8.15	5.94	31.3	35.99	4.55	3.35	30.3	4.09	3.15	26.0	1.38
Voigt average	A	#1065R006	7.89	6.96	12.6	9.15	4.55	4.35	4.3	4.46	4.07	9.2	0.39
	B	#1064R012	7.89	6.99	12.2	7.66	4.56	4.28	6.3	4.39	4.22	4.0	0.34
	C	#1066R013	8.65	6.12	34.2	33.61	4.89	3.71	27.5	4.43	3.48	23.9	1.41
Reuss average	A	#1065R006	6.61	6.02	9.3	10.96	3.59	3.35	7.0	3.50	3.13	11.3	0.37
	B	#1064R012	6.58	6.07	8.0	10.81	3.60	3.30	8.7	3.42	3.23	5.6	0.37
	C	#1066R013	7.61	5.74	28.0	39.56	4.17	2.95	34.2	3.74	2.78	29.5	1.37